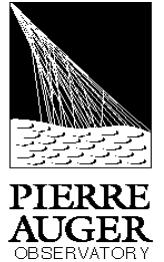


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# Annual Variation of the Atmosphere in Malargüe and the Implication on Auger FD Observables



setup in the Pampa Amarilla



Bianca Keilhauer



field PC in FD building  
Coihueco

Paris, May 2003



## data acquisition



- Auger Fluorescence Detector measures **longitudinal shower development**
- **Atmospheric parameter** affect the development and detection at every height

Knowledge of **atmospheric profiles** is required

**Radiosonde measurements** in each season are performed:

- 44 launches in total
- average reached altitude  $\sim$  20 km a.s.l. (maximum was 27.5 km a.s.l.)
- roughly every 20 m a set of data (**h, p, T, u, wind**)
- used DFM-97 GPS sondes ([www.graw.de](http://www.graw.de))
- accuracy:  $T < 0.2$  K
  - $p < 1.0$  hPa (range 200 hPa to 1080 hPa)
  - $< 0.5$  hPa (range 5 hPa to 200 hPa)
- $u < 5\%$

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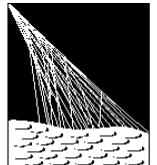
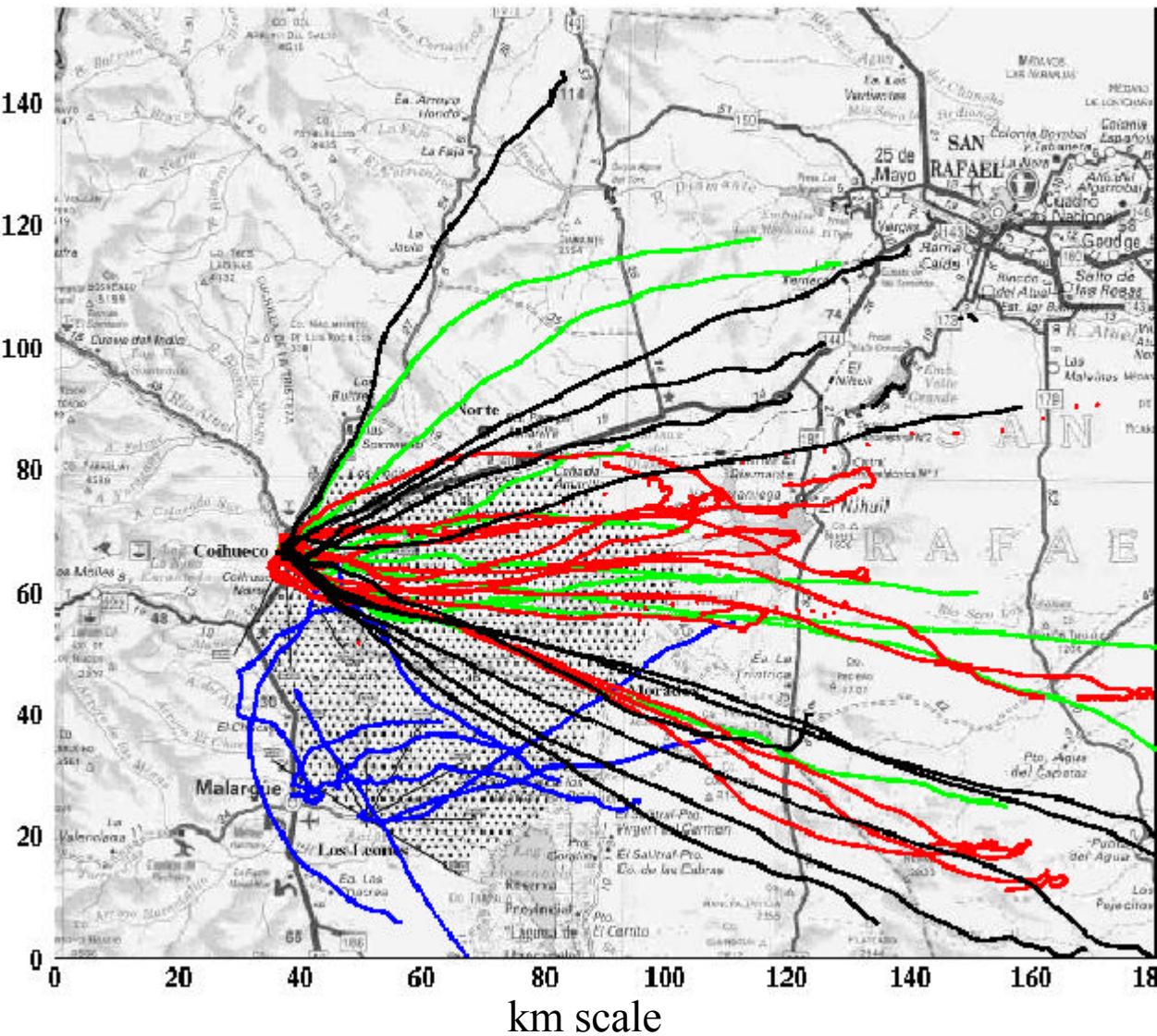
blue: winter,  
Aug. 2002

green: spring,  
Nov. 2002

red: summer,  
Jan./Feb. 2003

black: autumn,  
Apr./May 2003

## all balloon paths



PIERRE  
AUGER  
OBSERVATORY

Launches per  
campaign:

- winter – 9
- spring – 9
- summer – 15
- autumn - 11

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# important effects of atmospheric profiles

on the Auger FD shower data

1. atmospheric depth to geom. height

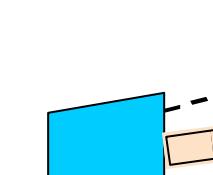
- $X = \int_0^h r(z) dz$

2. fluorescence light production

- fl. yield,  $\phi(p, T)$

3. fluorescence light transmission

- $t(p, T)$



transmission

telescope

fluorescence photons

height  
 $h$  (km)

$\frac{dE}{dx}$

$X$  to  $h$

Fl. Yield

p Fe

p Fe

atmosph. depth

$X_{\text{e}}^{\text{ag}} / \text{cm}^2 \text{ s}$

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## temperature variation

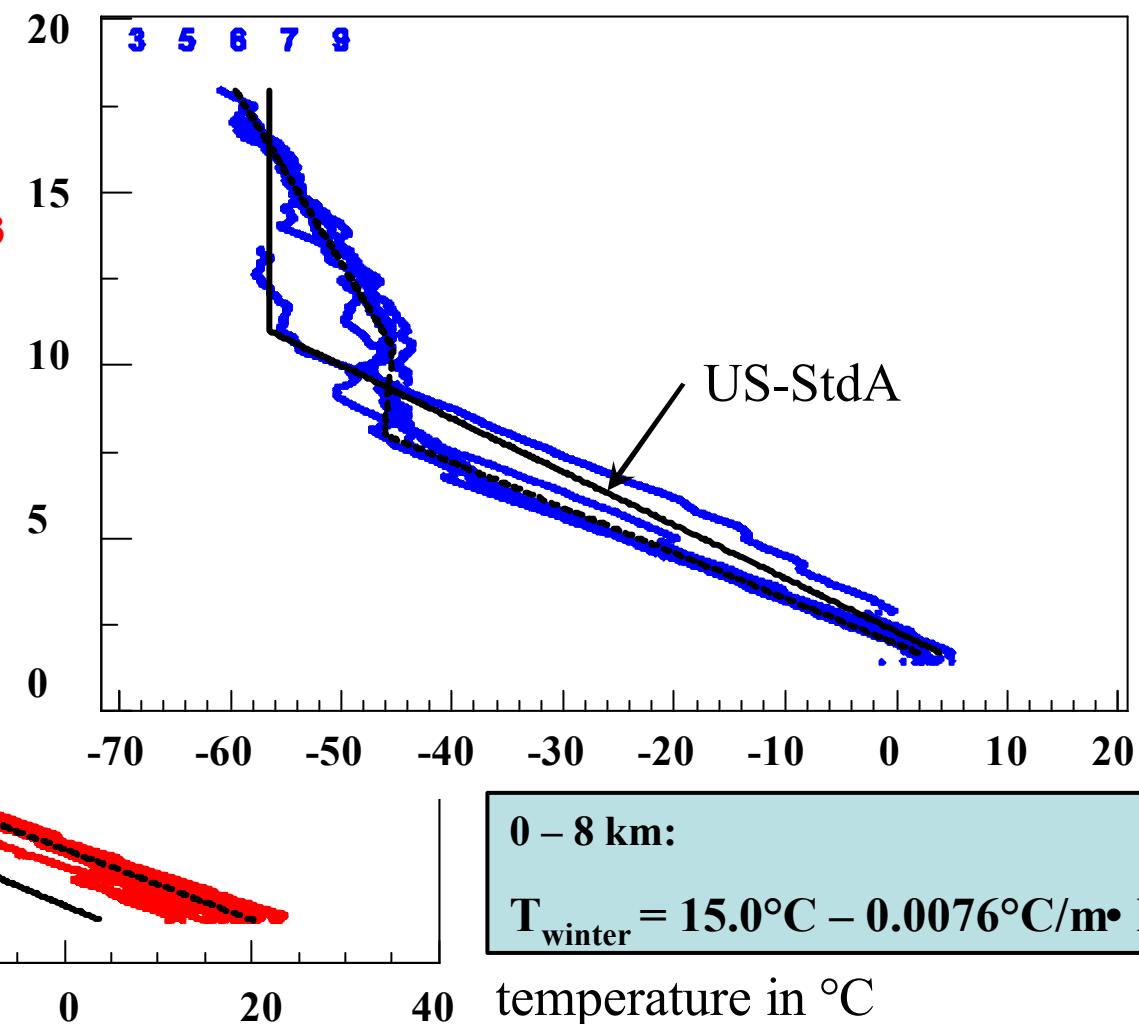
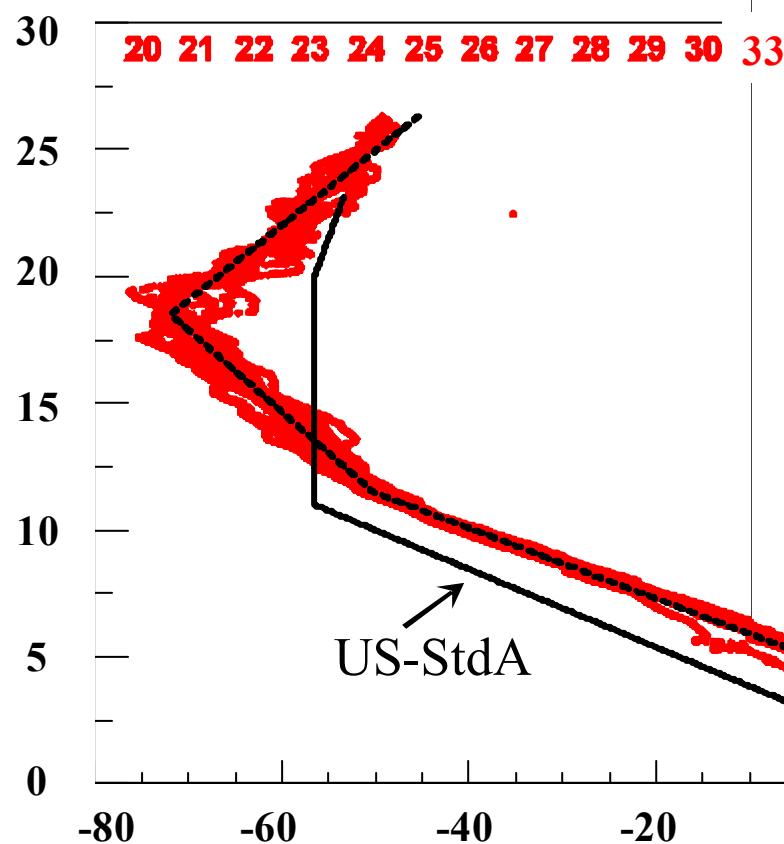
- summer and winter -



0 – 11.5 km:

$$T_{\text{summer}} = 32.5^{\circ}\text{C} - 0.0072^{\circ}\text{C}/\text{m} \cdot \text{h}$$

height in km a.s.l.



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## temperature variation

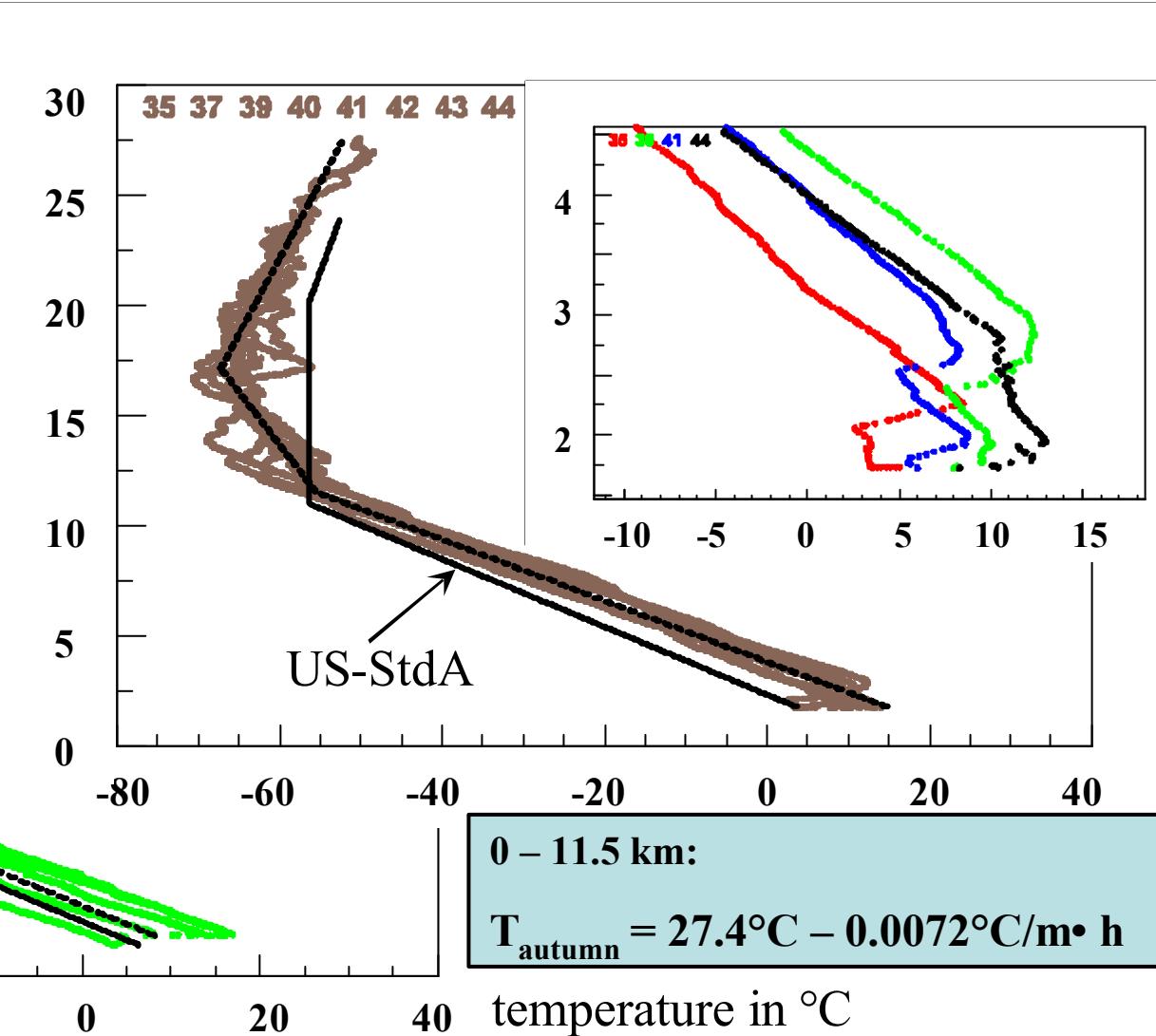
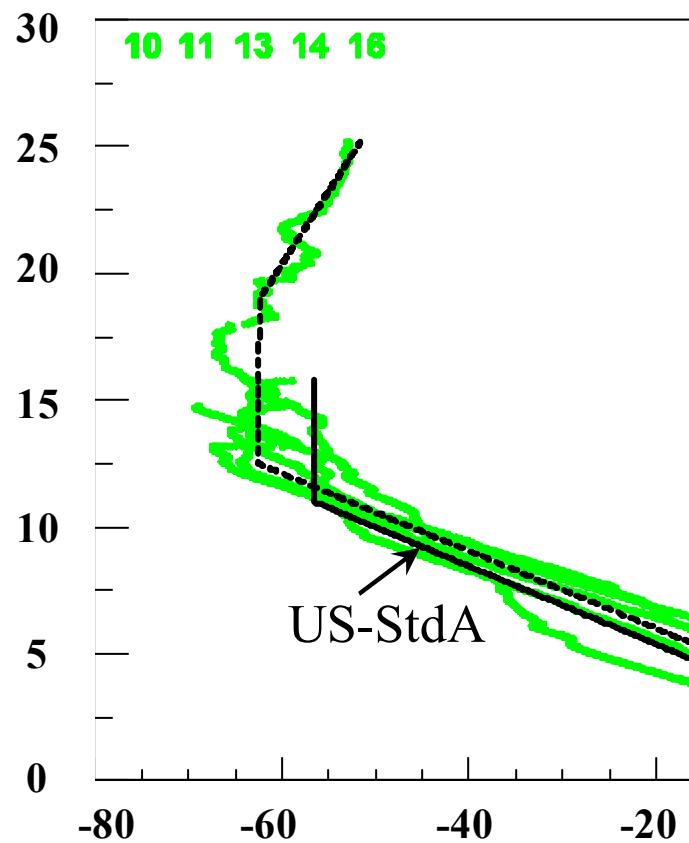
- spring and autumn -



0 – 12.5 km:

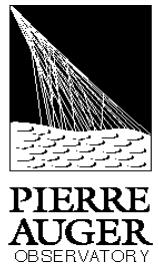
$$T_{\text{spring}} = 19.4^{\circ}\text{C} - 0.0066^{\circ}\text{C}/\text{m} \cdot \text{h}$$

height in km a.s.l.

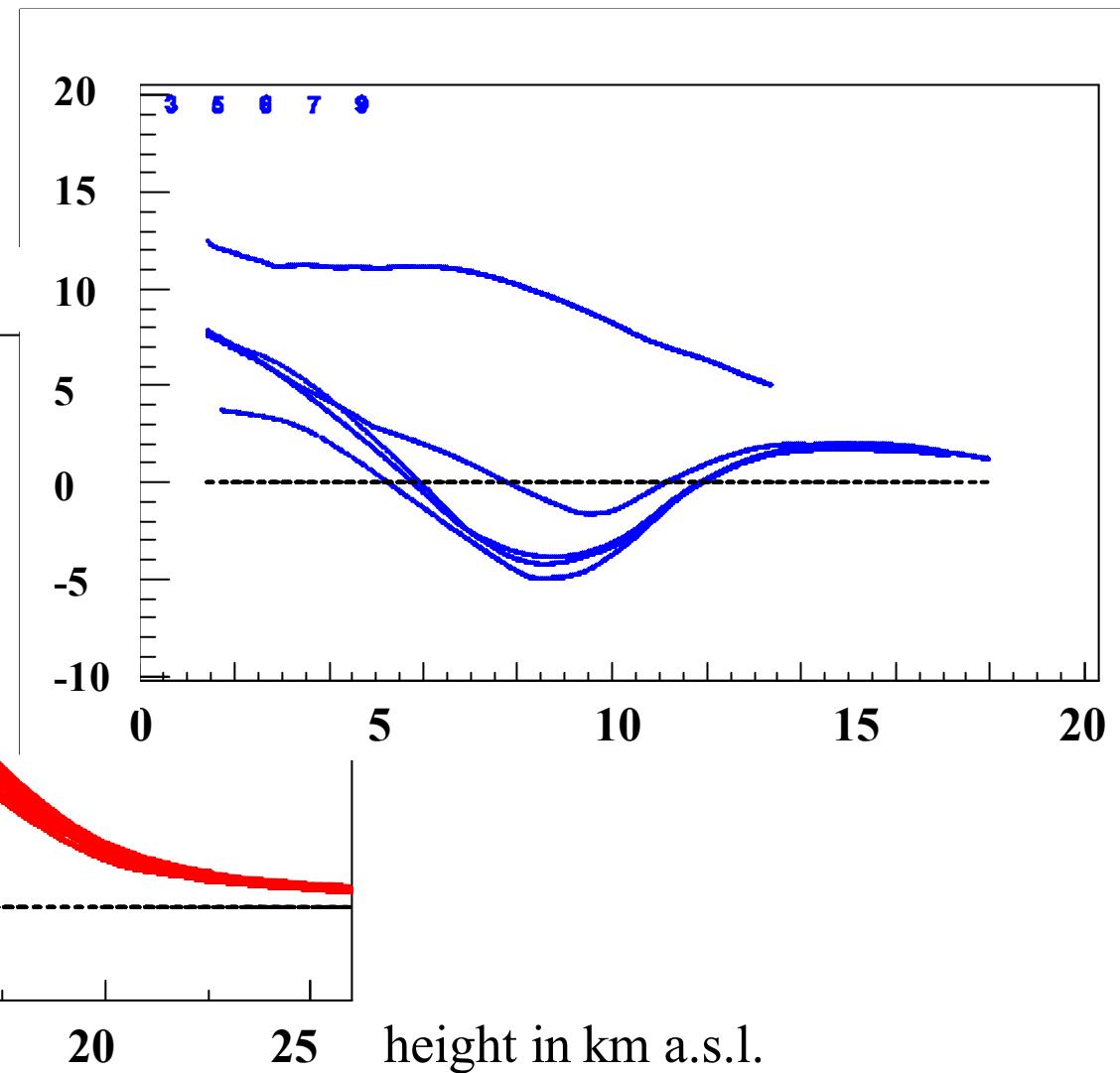
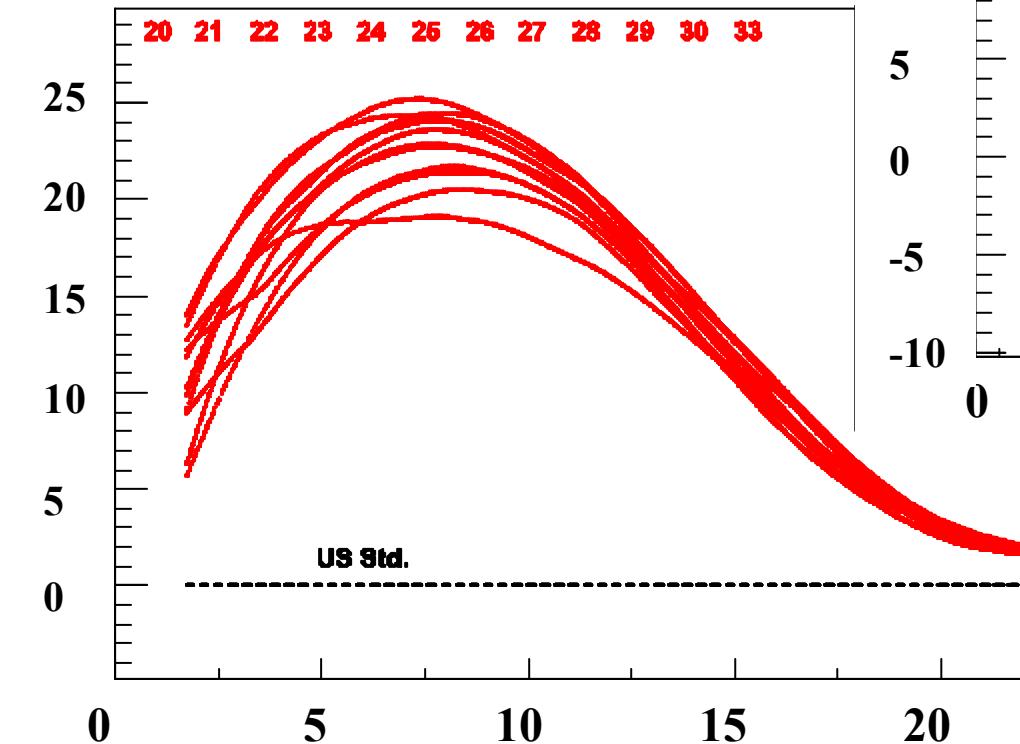


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## grammage variations - summer and winter -

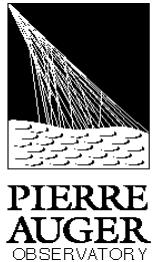


difference in atmos. depth in g/cm<sup>2</sup>

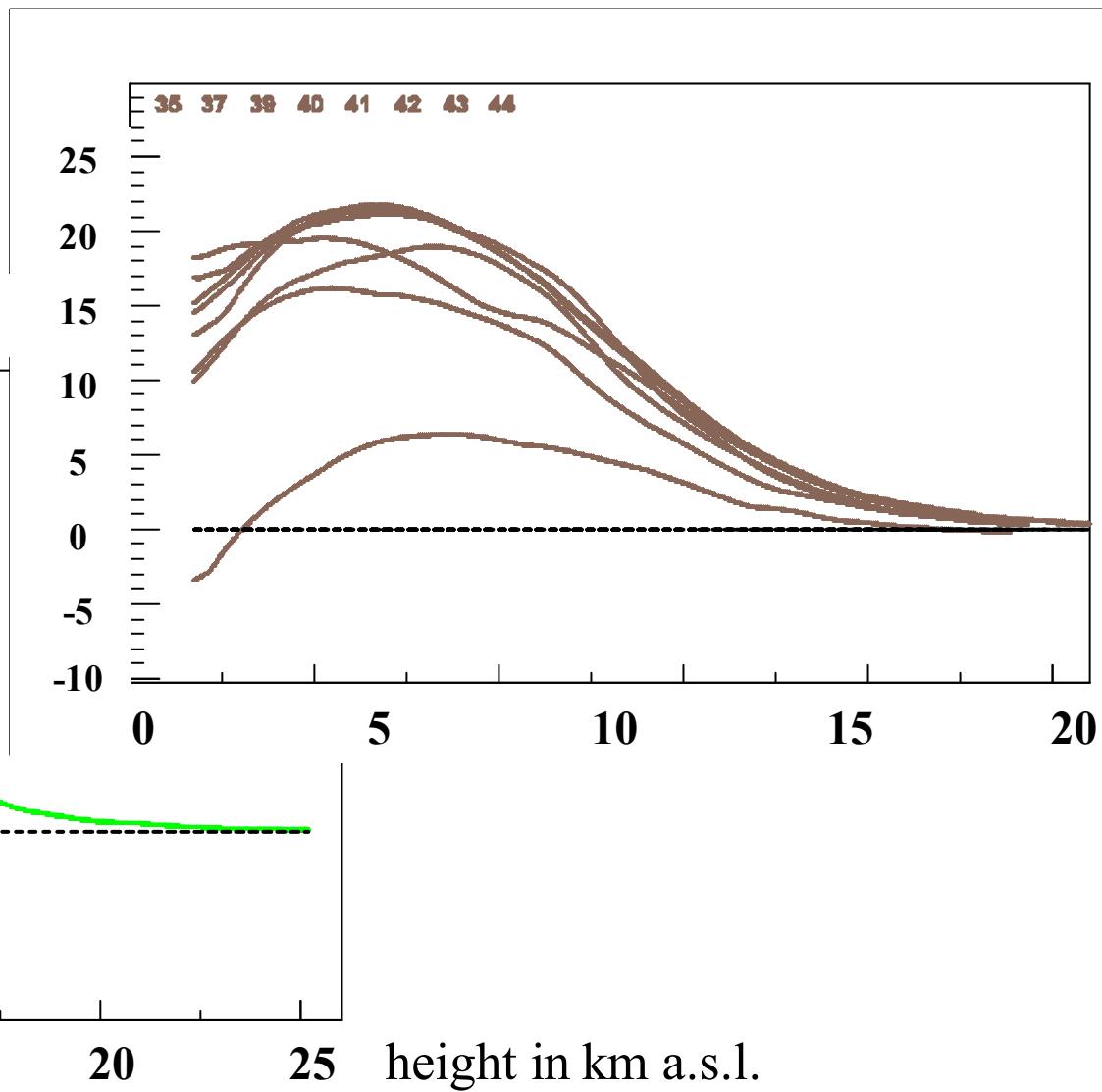
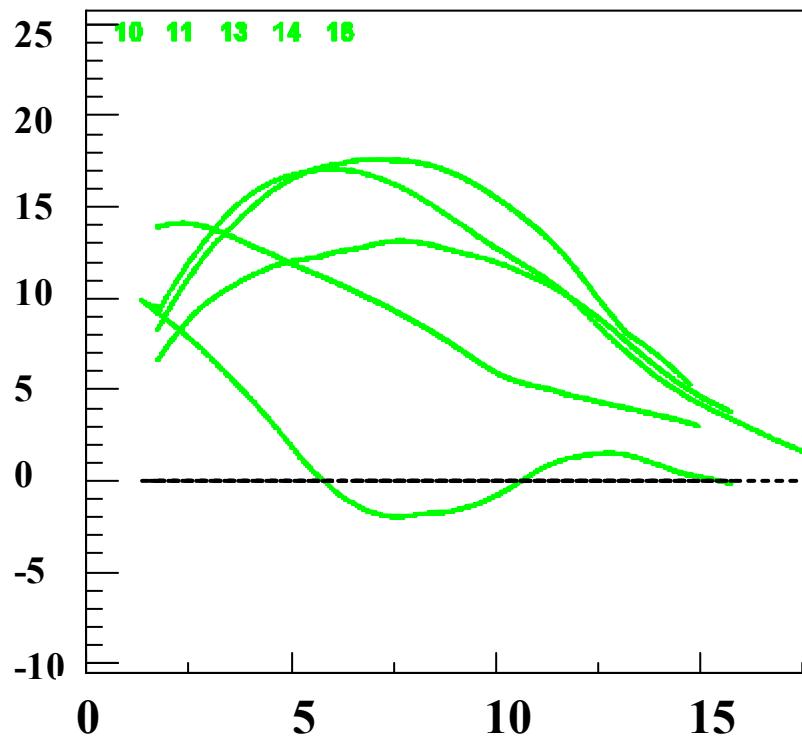


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## grammage variations - spring and autumn -

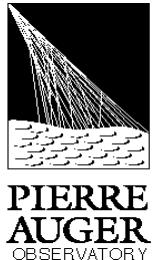


difference in atmos. depth in g/cm<sup>2</sup>

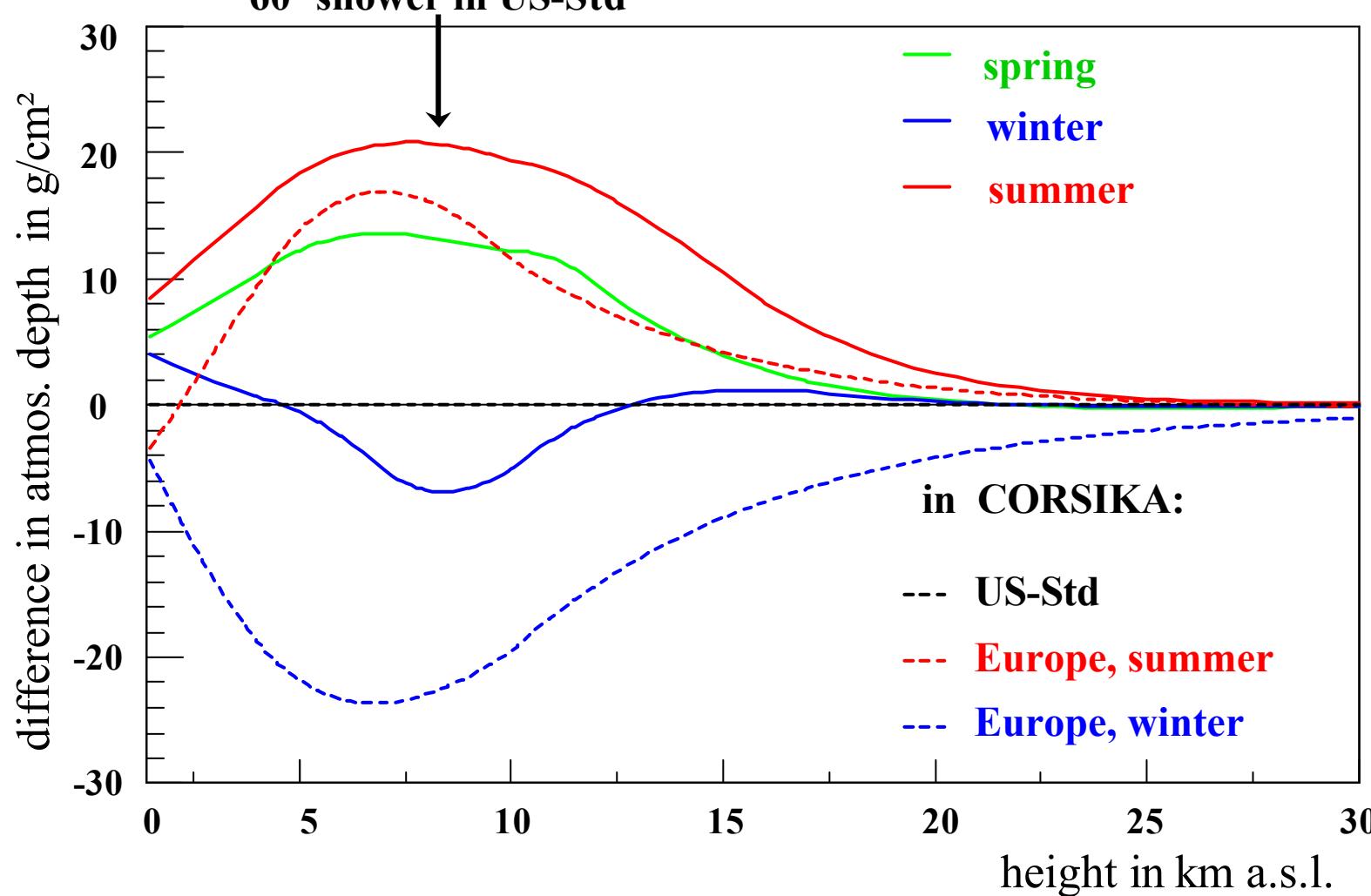




## typical grammage profiles



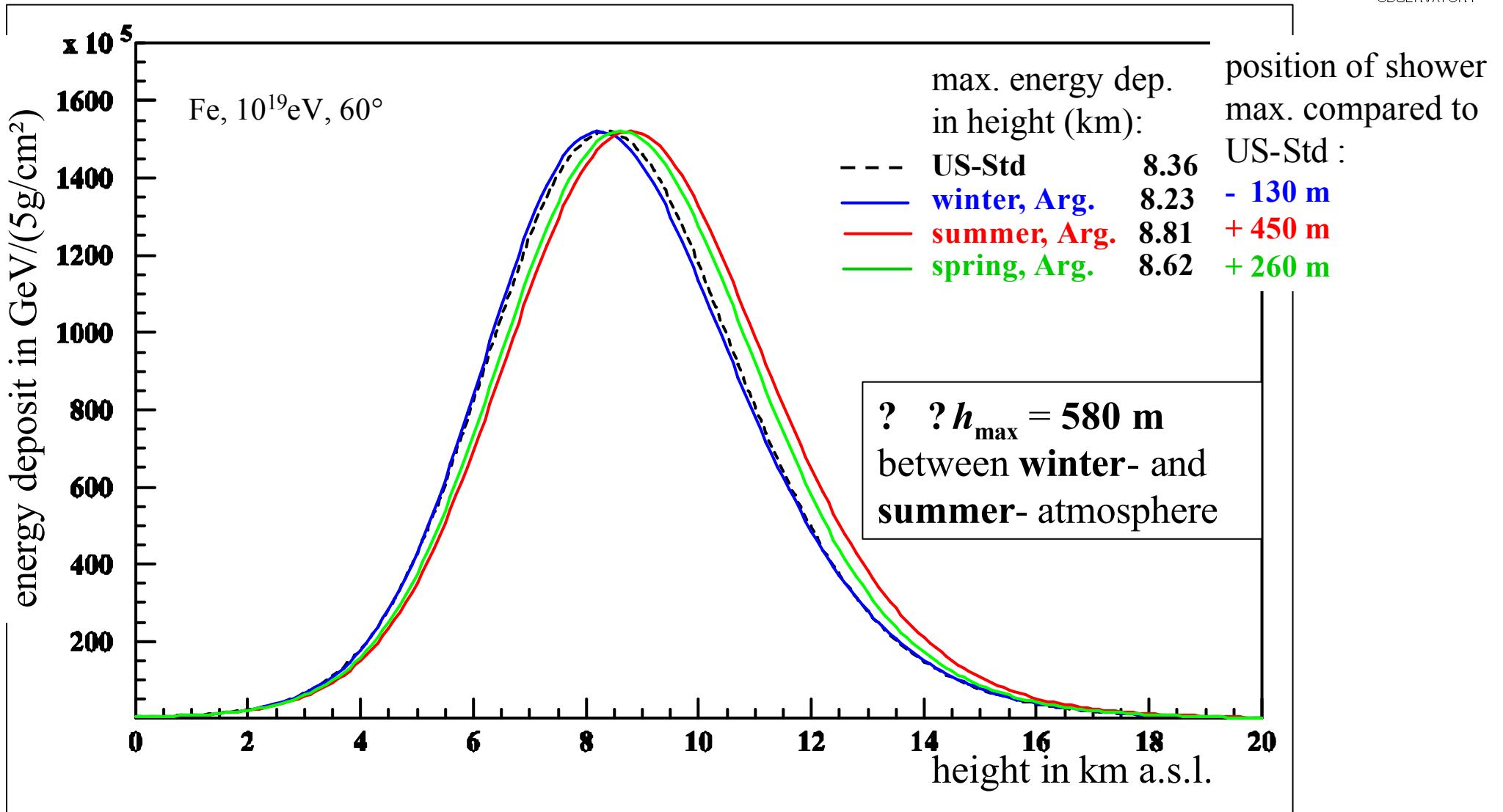
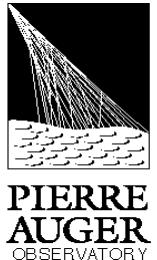
max. of Fe-ind.  $10^{19}$ eV,  
 $60^\circ$  shower in US-Std





# longitudinal shower development

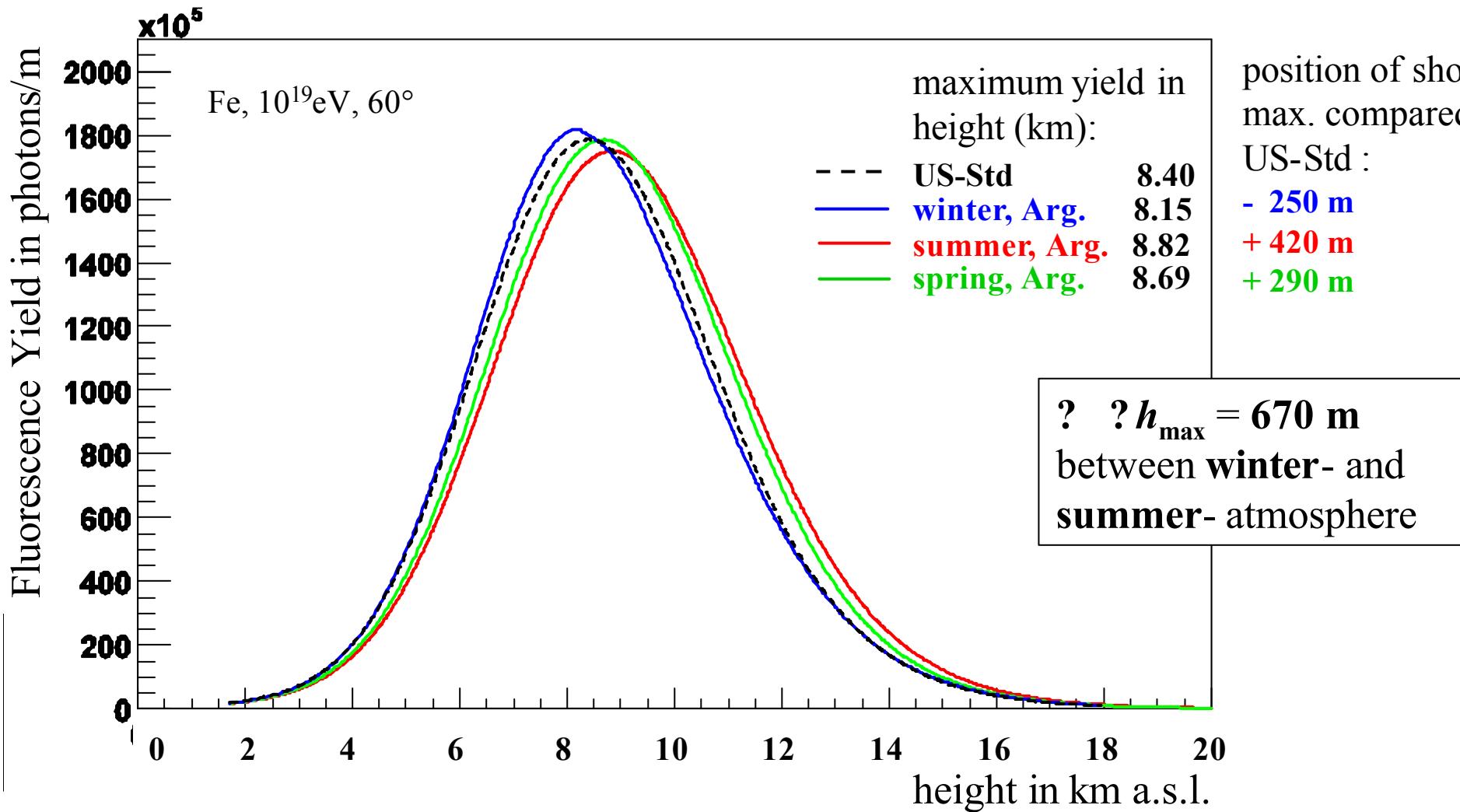
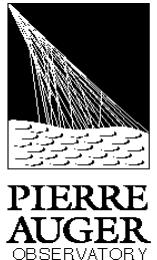
## - energy deposit -





# longitudinal shower development

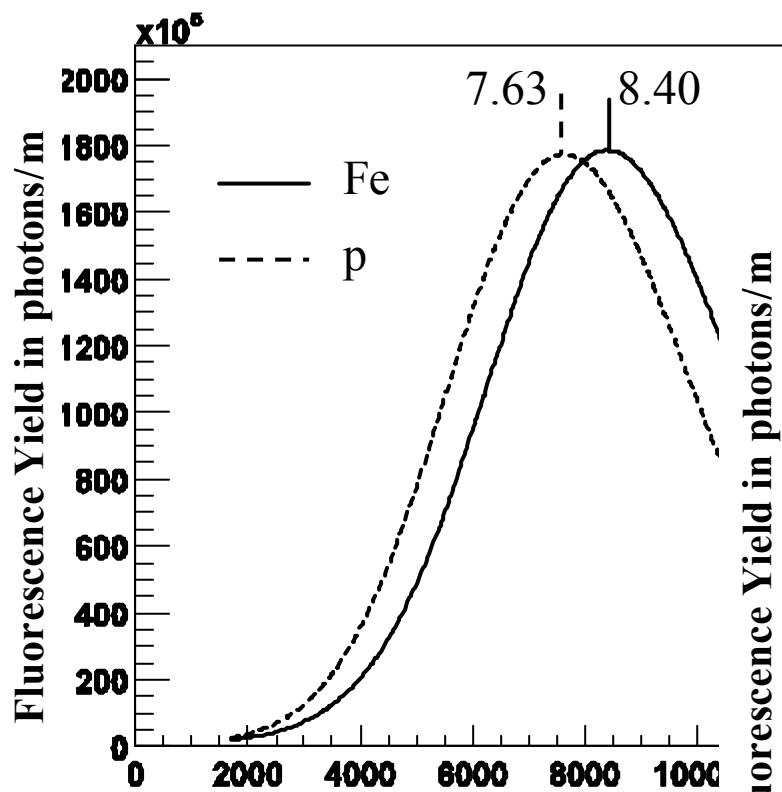
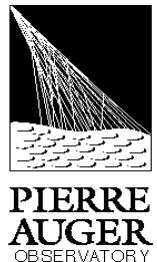
## - fluorescence yield -



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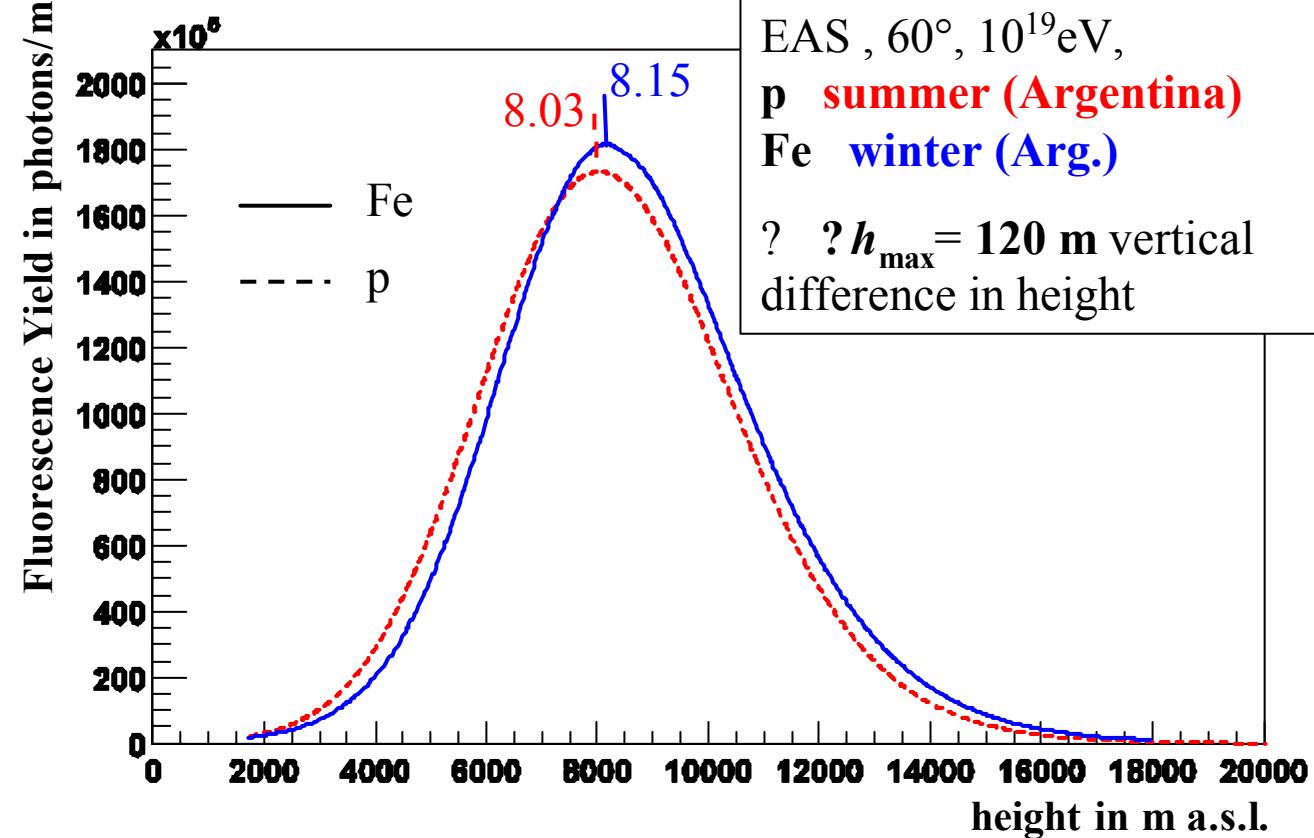
# position of shower maximum

## - fluorescence yield -



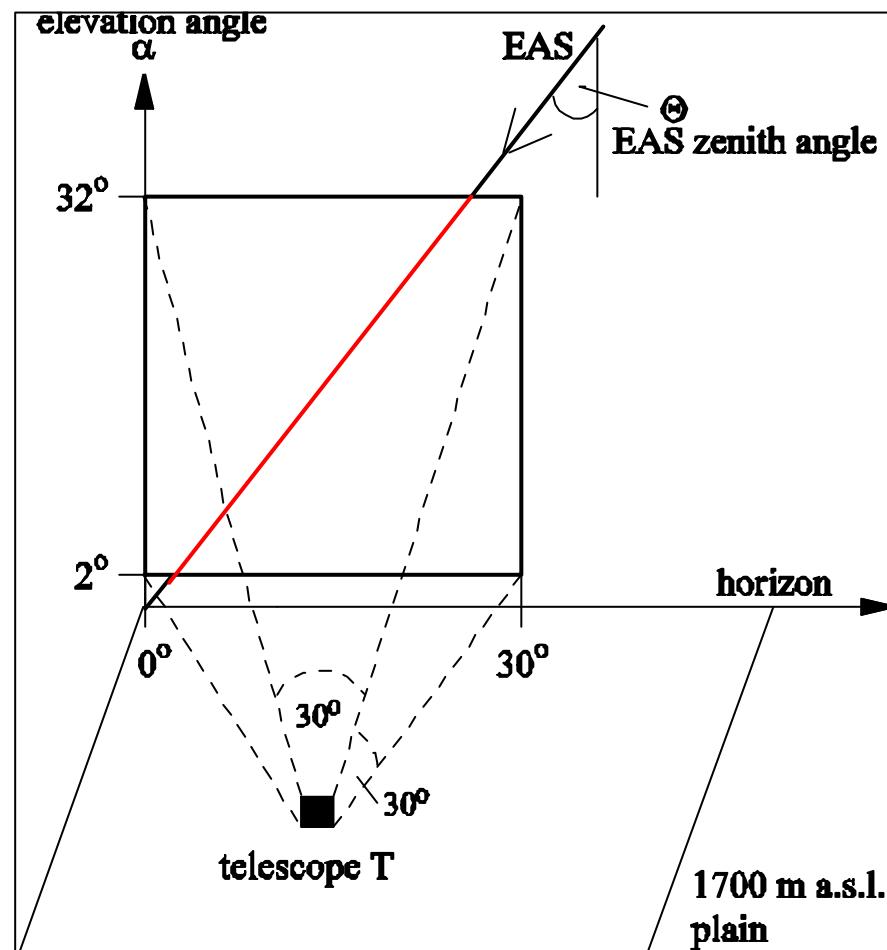
both EAS in US-Std,  $60^\circ$ ,  $10^{19}$ eV:

? ?  $h_{\max} = 770$  m vertical difference in height





# Rayleigh transmission - geometry



— visible part of  
the shower

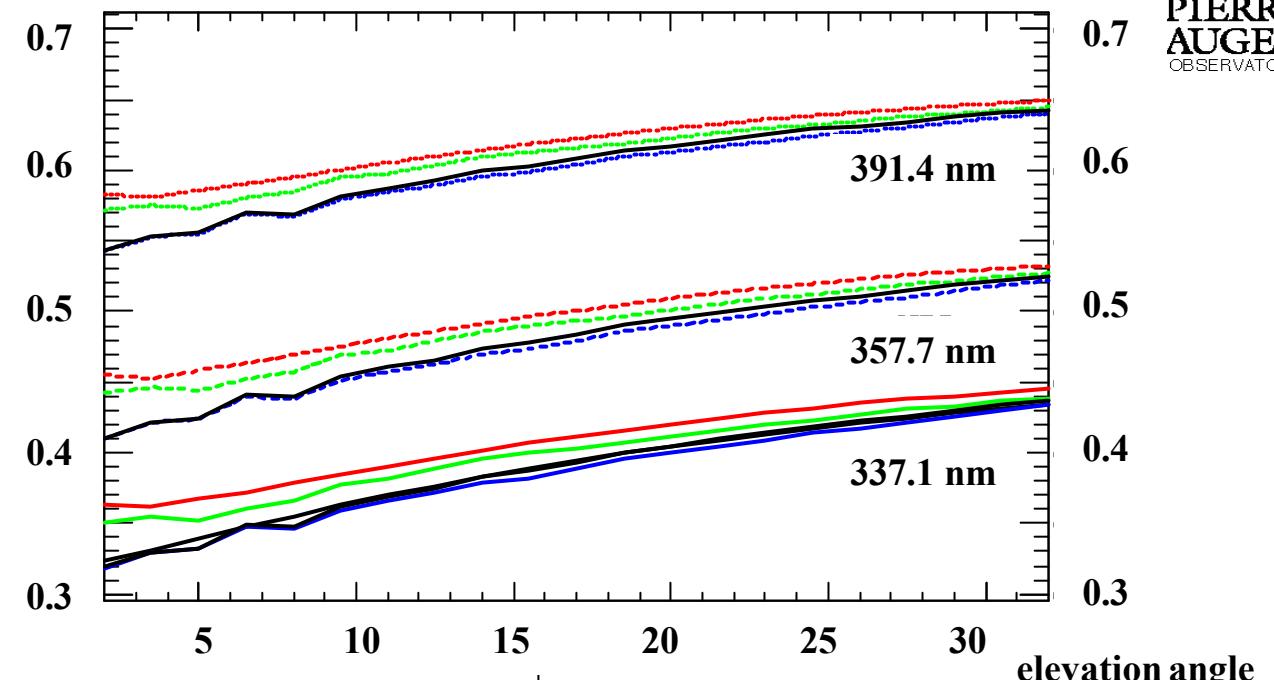
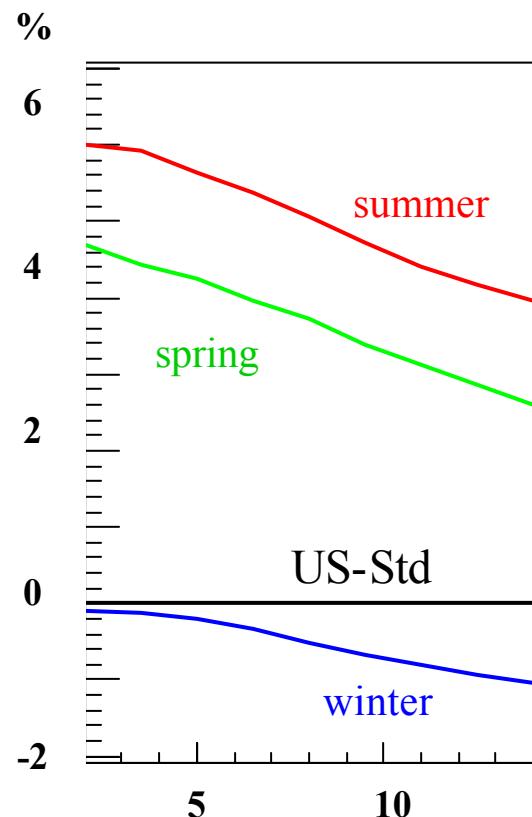
typical distance  
(shower, telescope):

~ 15 km

# Rayleigh transmission



realistic atmospheres  
- at three wavelengths



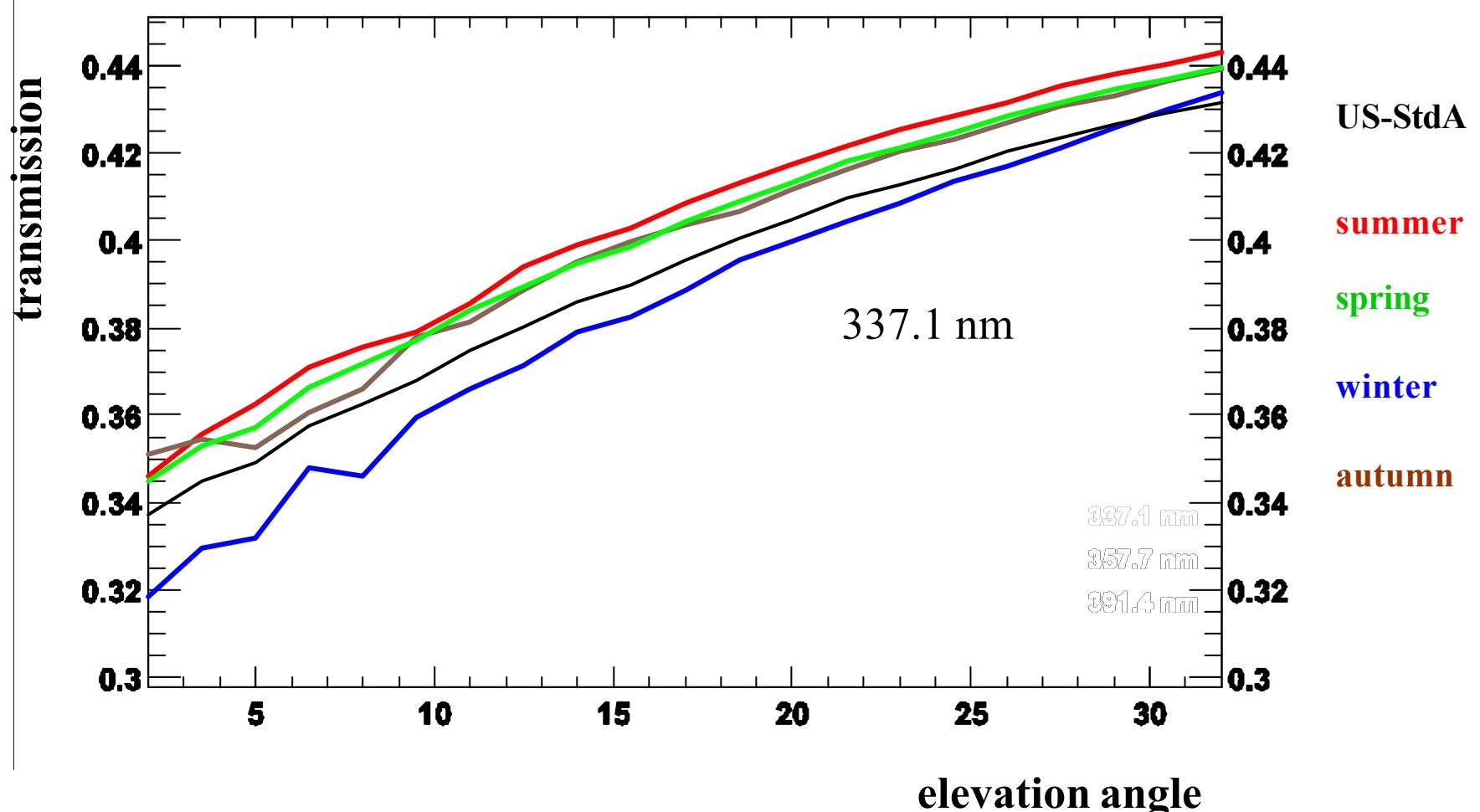
differences in Rayleigh  
transition at 337.1 nm

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## Rayleigh transmission

with correction for ground parameter



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## summary



- **three important effects of the atmospheric profiles :**
  1. longitudinal shower development according to geometrical height  
(?  $h_{max,Fe}$  (summer ? winter) up to 580 m)
  2. fluorescence light production  
(? *Fl.Yield* from measured data  $\sim 4\%$ , with constant  $E_{dep}$  ;  
?  $h_{max,Fe}$  (summer ? winter) up to 670 m)
  3. fluorescence light transmission to the telescope (Rayleigh)

good data on the atmospheric profiles is necessary for the accurate reconstruction of EAS

$$(\text{? } h_{max} \text{ up to } (30 \text{ g/cm}^2)/\cos T \quad \text{? } h_{max} (\text{p to Fe}) \sim 80 \text{ g/cm}^2)$$

- **first radiosonde data from Malargüe, Argentina :**
  1. large differences from the US Standard atmosphere
  2. seasonal variations shifted compared to Europe data